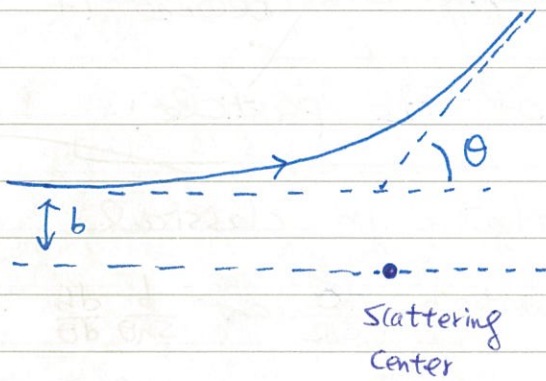


Motivation: One of the best ways to understand the structure of particles and interactions between them is to scatter them off each other: X-ray, photo-emission, ...
In fact, the images that we observe everyday are results of scattering.

Here, let us start from classical scattering theory first.

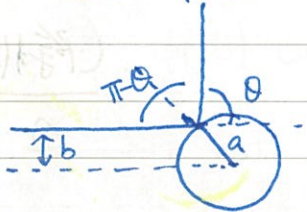
Classical scattering theory (with a fixed scattering center)



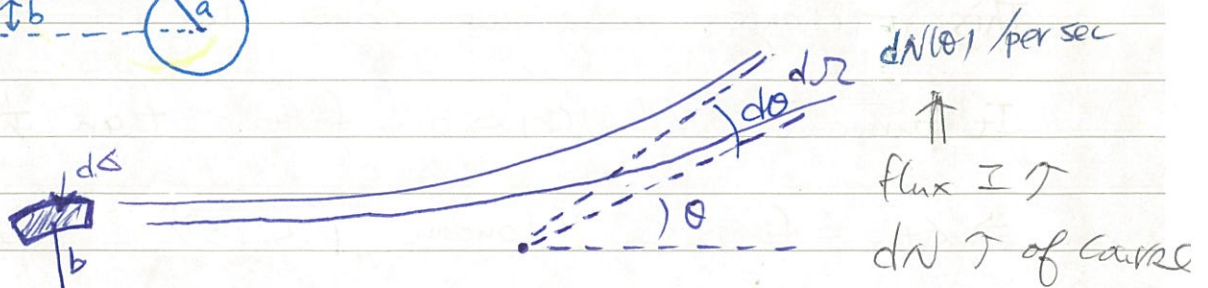
assume: recoil is negligible
azimuthally symmetrical.

b: describes "accuracy of aiming"
→ impact parameter
theta: Scattering angle.

example: hard-sphere scattering



$$b = a \sin\left(\frac{\pi}{2} - \frac{\theta}{2}\right) = a \cos\frac{\theta}{2}$$



This figure shows

of particles with impact parameter $\in (b, b+db)$ must correspond to # of particles scattered into $d\Omega$

$$\Delta(\theta) = \frac{d\sigma}{d\Omega} = \text{differential cross-section}$$

$$d\sigma = \frac{dN(\theta)}{I} = \text{function of } \theta \equiv \Delta(\theta) d\Omega$$

$$(dN(\theta)) = 2\pi b db \cdot I \neq I db$$

From $I \rightarrow \langle \theta \rangle$ just changes of variables

from $b \rightarrow \theta$ to $\sin \theta \rightarrow \theta$

Date

2

$$\therefore d\sigma = b(b) d\phi$$

$$dR = \sin \theta d\theta d\phi$$

$$\therefore \frac{d\sigma}{dR} = \frac{b}{\sin \theta} \frac{db}{d\theta}$$

$$\left(\because I \propto b db = -I d\sigma \right)$$

$$= -I \langle \theta \rangle dR$$

using # of particles $\propto \sin \theta d\theta$
not changed! $I \cdot dA = I \langle \theta \rangle dR$

Typically, θ is a decreasing function of b

$\therefore \frac{db}{d\theta} < 0$. \therefore one chooses positive value by

setting $\frac{d\sigma}{dR} = \frac{b}{\sin \theta} \left| \frac{db}{d\theta} \right|$

total cross-section

$$\sigma \equiv \int \left(\frac{d\sigma}{dR} \right) \cdot dR$$

$$\sigma = \int \langle \theta \rangle dR$$

= trace all area

unit = area, meaning: track down effective areas into which when incident particles move will be scattered!

↳ see over

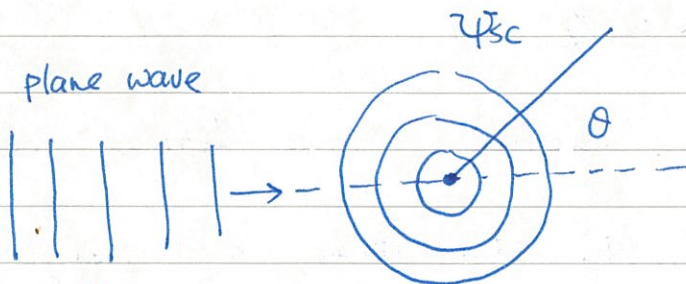
example: hard-sphere scattering: $\frac{d\sigma}{dR} = \frac{a^2}{4}$

$$b = a \cos \theta/2$$

$$\sigma = \pi a^2$$

Quantum Scattering

general elastic theory:

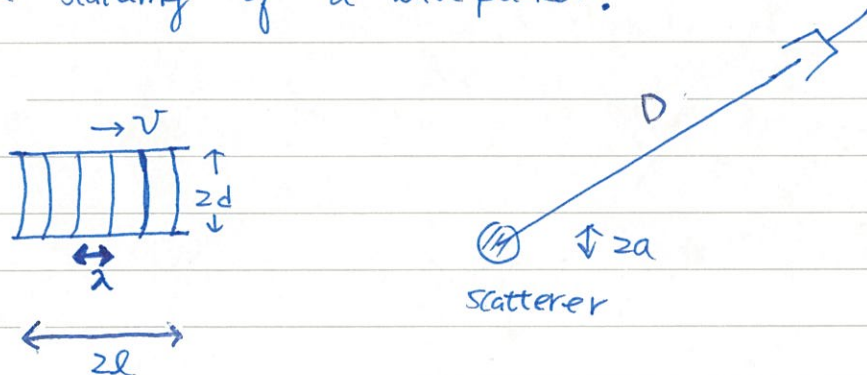


↑
most general: wave packet

$$\psi \rightarrow \left(\frac{1}{\sqrt{2\pi}} \right)^3 e^{ik \cdot x} + \psi_{s.c.}$$

↑
original plane wave

Rigorously speaking, in reality, one must deal with the scattering of a wavepacket:



However, if $a \ll d, l$, the collision phenomenon does not depend critically on the particular form of the wave packet.

Furthermore, if $\lambda \ll d$ & $\lambda \ll l$, the incident wave are well approximated by a plane wave (\therefore

$$\Delta x \sim l, \Delta y \sim d, \Delta p_x \sim \frac{\hbar}{l}, \Delta p_y \sim \frac{\hbar}{d}$$

$$\Delta k \propto \frac{\Delta \lambda}{\lambda} \cdot \frac{1}{\lambda}$$

$$\frac{\Delta \lambda}{\lambda} \sim \frac{1}{l} \ll 1$$

$$\Delta \delta_0 = d \text{ or } l$$

$$\Delta \delta = \left[(\Delta \delta_0)^2 + \left(\frac{\Delta p_x}{m} \right)^2 \right]^{\frac{1}{2}}$$

Further consideration of spreading leads to $\sqrt{\lambda D} \ll d, l$ (to avoid effects due to spreading)

See Messiah's Quantum Mechanics Chapter X for more details.

$$\frac{\Delta p_x t}{m} \sim \frac{\hbar t}{2m \Delta \delta_0} = \frac{100 \hbar}{2000 m v}$$

$$\Delta \delta_0 \Delta \delta = \hbar/2 = \frac{D \hbar}{2 \Delta \delta_0}$$

\therefore We can simply solve the problem for a single.

Wave vector k : $\psi_k(r) e^{-\frac{i}{\hbar} E_k t}$, elastic: $E_k = \frac{\hbar^2 k^2}{2m}$

$$\Delta \delta = \Delta \delta_0 \left(1 + \frac{1}{4} \left(\frac{D \hbar}{\Delta \delta_0^2} \right)^2 \right)^{\frac{1}{2}}$$

$$\frac{-\hbar^2}{2m} \nabla^2 \psi_k + V \psi_k = E_k \psi_k$$

The problem then becomes time-independent.

(Note that if one really wants to treat wavepackets, suppose $V(r) \rightarrow 0$ as $r \rightarrow \infty$, one can form packets from the single k solutions)

$\psi_k(r) \rightarrow$ free particles

Recall that radial part = $\frac{u_{\ell}(r)}{r}$ which

$$\text{satisfies } \frac{d^2 u_{\ell}}{dr^2} + \frac{2m}{\hbar^2} \left[E - \underbrace{V(r)}_{\equiv k^2 + U(r)} \right] u_{\ell} = 0$$

$$u_{\ell}(0) = 0$$

$$k^2 + U(r)$$

where $U(r) = \frac{-2m}{\hbar^2} V(r) - \frac{\ell(\ell+1)}{r^2}$. (ii) Suppose $V(r)$ slows than $\frac{1}{r^2}$

We can drop ℓ . Try $u_{\ell} = f e^{\pm ikr}$

$$\Rightarrow f'' \pm 2ik f' + U(r) f = 0$$

since we expect

$$\begin{aligned} & \text{if } f' \sim Uf \quad \downarrow \frac{1}{\sqrt{U}} \quad \uparrow \frac{1}{\sqrt{U}} \\ & f'' = Uf' + Uf' \\ & \text{if } |f''| \ll |2ik f'| \\ & \uparrow \quad \uparrow \\ & \frac{1}{\sqrt{U}} \text{ or } \frac{1}{\sqrt{2U}} \quad U(r) f' \\ & U' \quad (U)^2 \quad \frac{1}{r^2} \\ & \text{self-const.} \end{aligned}$$

$$\therefore f(r) = f(r_0) \exp\left[\frac{\pm}{2ik} \int_{r_0}^r U(r') dr'\right]$$

$f(r) \rightarrow \text{const}$ only when $U(r) \cdot r \rightarrow 0$

The above derivation counts on the spherical symmetry. \rightarrow jump to Schwinger-Lipp. eq.

We will have more justification later.

We have seen, for free particles (ii) $U(r) \rightarrow 0$ faster than

$$\psi(r \rightarrow \infty) = \sum_{\ell m} [A_{\ell} j_{\ell}(kr) + B_{\ell} h_{\ell}(kr)] Y_{\ell}^m(\Omega) \quad \frac{1}{r^2}$$

$$\therefore j_e(kr) \rightarrow \frac{\sin(kr - \ell\pi/2)}{kr}$$

$$n_e(kr) \rightarrow \frac{-\cos(kr - \ell\pi/2)}{kr}$$

$$\left(\frac{e^{-ikr}}{r} (-1)^\ell \right)$$

$$\frac{A_e}{B_e} = -i \Rightarrow \text{get } e^{ikr} \text{ outgoing}$$

$$\therefore \psi(\vec{r}) \rightarrow e^{i\vec{k}\cdot\vec{r}} + \frac{e^{ikr}}{kr} \sum_{\ell m} (-i)^\ell (B_\ell) Y_\ell^m(\Omega)$$

$$\equiv e^{i\vec{k}\cdot\vec{r}} + \frac{e^{ikr}}{r} f(\theta, \phi) \dots \textcircled{1}$$

↑

fact that the scattered wave takes the form: spherical.

$\frac{e^{ikr}}{r} f(\theta, \phi)$ is not surprising and is expected.

since the probability is conserved: what is flowing out radius r should also flow out $r' > r$, otherwise, something will be created between r and r' !

f is known as scattering amplitude.

We shall show that

$$\frac{d\sigma}{d\Omega} = |f(\theta, \phi)|^2 \dots \textcircled{2}$$

so $|f|^2$ has the meaning of differential crosssection.

To show eq. ②, we have to calculate the # of particles that flow into $d\Omega$, i.e., the probability current.

$$|\vec{J}_{inc}| = \left| \frac{\hbar}{2mi} (e^{-i\vec{k}\cdot\vec{r}} \nabla e^{i\vec{k}\cdot\vec{r}} - e^{i\vec{k}\cdot\vec{r}} \nabla e^{-i\vec{k}\cdot\vec{r}}) \right| \cdot \frac{1}{V}$$

← normalization constant of $e^{i\vec{k}\cdot\vec{r}}$

$$= \frac{\hbar k}{m} \cdot \frac{1}{V} \left(\frac{1}{\sqrt{V}} e^{i\vec{k}\cdot\vec{r}} \right) \quad (\text{particles / sec area})$$

Same normalization as in eq. (3)

$$|\vec{J}_{sc}| = \left| \frac{\hbar}{2mi} \left(\frac{e^{-i\vec{k}\cdot\vec{r}}}{r} f^* \nabla \frac{e^{i\vec{k}\cdot\vec{r}}}{r} f - \frac{e^{i\vec{k}\cdot\vec{r}}}{r} f \nabla \frac{e^{-i\vec{k}\cdot\vec{r}}}{r} f^* \right) \right| \cdot \frac{1}{V}$$

← as in eq. (3)

$$\nabla = \hat{r} \frac{\partial}{\partial r} + \hat{\theta} \frac{1}{r} \frac{\partial}{\partial \theta} + \hat{\phi} \frac{1}{r \sin \theta} \frac{\partial}{\partial \phi}$$

$O(1/r^2)$ when acting on $(\frac{e^{i\vec{k}\cdot\vec{r}}}{r} f)$

$$\frac{\partial}{\partial r} \frac{e^{i\vec{k}\cdot\vec{r}}}{r} f = ik \frac{e^{i\vec{k}\cdot\vec{r}}}{r} f + O(1/r^2)$$

f is defined under the normalization form in eq. (3)

$$\therefore \vec{J}_{sc} = \frac{\hat{r}}{r^2} |f(\theta, \phi)|^2 \frac{\hbar k}{m} \left(\frac{1}{V} \right) + O(1/r^3)$$

← eq. (3)

particles / per sec through $d\Omega = \vec{J}_{sc} \cdot \hat{r} r^2 d\Omega$

$$= |f|^2 \frac{\hbar k}{m} d\Omega \cdot \left(\frac{1}{V} \right)$$

$$\therefore \frac{|f|^2 \frac{\hbar k}{m} d\Omega}{\frac{\hbar k}{m}} = \frac{dG}{d\Omega} \cdot d\Omega$$

$$\frac{dG}{d\Omega} = |f(\theta, \phi)|^2$$

Note that rigorously speaking, there are interference terms between $e^{i\vec{k}\cdot\vec{r}}$ & $\frac{e^{i\vec{k}\cdot\vec{r}}}{r} f(\theta, \phi)$ when calculating the current. However, since in real experiments, it



Many scattering centers

(often encountered in Condensed Matter physics)

or $N_S(r-r_i)$ (classically)

$$\psi_S(r) = \sum_i \psi_S(r-r_i) \quad (\text{assuming no spin})$$

$$I \propto \langle |\psi_S(r)|^2 \rangle$$

↑

average over positions of r_i

$$= \sum_{i,j} \langle \psi_S(r-r_i) \psi_S(r-r_j) \rangle$$

$$= \int dr_1 \int dr_2 \langle \sum_i \delta(r_1-r_i) \psi_S(r-r_1) \sum_j \delta(r_2-r_j) \psi_S(r-r_2) \rangle$$

$$= \int dr_1 \int dr_2 \underbrace{\langle n(r_1) n(r_2) \rangle}_{\text{correlation function}} \psi_S(r-r_1) \psi_S(r-r_2)$$

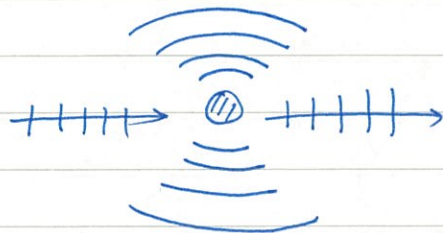
correlation function

If spins are involved

$$\Rightarrow \langle \vec{S}(r_1) \cdot \vec{S}(r_2) \rangle$$

∴ Scattering type expts. measure correlations.

occurs in the following way



\therefore We may drop the interference terms.

The Lippmann-Schwinger equation

Let us write $H = H_0 + V$, $H_0 = \frac{p^2}{2m}$

$$V = V(\mathbf{r})$$

We would like to develop ^{the} perturbation theory around

$$H_0 |\phi\rangle = E |\phi\rangle$$

For elastic scatterings, we need to solve

$$(H_0 + V) |\psi\rangle = E |\psi\rangle \quad \text{for the same } E.$$

Since $(E - H_0) |\psi\rangle = V |\psi\rangle$ & $(E - H_0) |\phi\rangle = 0$,

We may try to write $|\psi\rangle = |\phi\rangle + \frac{V}{E - H_0} |\psi\rangle$

This is not quite right as we have shown in

5/8/96, 2: the correct way to invert $E - H_0$ is

$$|\psi^\pm\rangle = |\phi\rangle + \frac{1}{E - H_0 \pm i\epsilon} V |\psi^\pm\rangle$$

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As we have also shown, in real space, this equation leads to

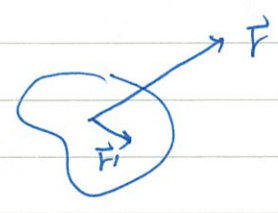
integral converges only when $|r'| \rightarrow 0$ as $r \rightarrow \infty$

$$\psi^\pm(\vec{r}) = \phi_p(\vec{r}) - \frac{2m}{\hbar^2} \int d^3r' \frac{e^{\pm i k |\vec{r} - \vec{r}'|}}{4\pi |\vec{r} - \vec{r}'|} V(\vec{r}') \psi^\pm(\vec{r}') \quad k = |\vec{p}|$$

$\therefore |\vec{r} - \vec{r}'| \approx r$ as $r \rightarrow \infty$

For a finite range potential, the region that $V(\vec{r}')$ contributes is limited in space.

\therefore We can approximate $r \gg r'$ in the integral:



$$\begin{aligned} |\vec{r} - \vec{r}'| &= r \sqrt{1 - \frac{2r'}{r} \hat{r} \cdot \hat{r}' + \left(\frac{r'}{r}\right)^2} \\ &\approx r \left(1 - \frac{2r'}{r} \hat{r} \cdot \hat{r}' + o\left(\frac{r'}{r}\right)^2\right) \\ &= r - \vec{r} \cdot \vec{r}' \end{aligned}$$

If we define $\vec{k}' \equiv k \hat{r}$, which is the wave vector for the outgoing particle, we have

$$e^{\pm i k |\vec{r} - \vec{r}'|} = e^{\pm i k r} e^{\mp i \vec{k}' \cdot \vec{r}'}$$

$$\begin{aligned} \therefore \psi^+(\vec{r}) &\rightarrow \phi_p(\vec{r}) - \frac{1}{4\pi} \frac{2m}{\hbar^2} \frac{e^{i k r}}{r} \int d^3r' e^{-i \vec{k}' \cdot \vec{r}'} V(\vec{r}') \psi^+(\vec{r}') \\ &= \frac{1}{(2\pi)^{3/2}} \left[e^{i \vec{k} \cdot \vec{r}} + \frac{e^{i k r}}{r} f(\vec{k}', \vec{R}) \right] \end{aligned}$$

*$\phi_p = \int d^3k |k\rangle \langle k| \phi_p$
 $+ \int d^3k |l\rangle \langle l| \phi_p$
 \dots*

\rightarrow perturbative view if box normalization is used. $\frac{1}{(2\pi)^{3/2}} \rightarrow \sqrt{\Omega}$

$$f(\vec{k}', \vec{R}) = \frac{-1}{4\pi} \frac{2m}{\hbar^2} (2\pi)^3 \int d^3r' \frac{e^{-i \vec{k}' \cdot \vec{r}'}}{(2\pi)^{3/2}} V(\vec{r}') \langle \vec{r}' | \psi^+ \rangle$$

$$= \frac{-1}{4\pi} (2\pi)^3 \frac{2m}{\hbar^2} \langle \vec{k}' | V | \psi^+ \rangle = f(\theta, \phi) \Rightarrow \text{back to interpret } f$$

Now, ③ is correct even $V(\vec{r}') \neq V(r)!$

3/4/1991 * Born Approximation

One can iterate eq. (3) and get the

1st order Born Approximation

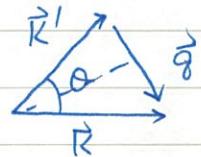
$$f^{(1)}(\vec{k}, \vec{k}') = -\frac{1}{4\pi} \frac{2m}{\hbar^2} \int d^3\vec{r}' e^{i(\vec{k}-\vec{k}') \cdot \vec{r}'} V(\vec{r}')$$

$$\therefore \frac{d\sigma}{d\Omega} = |f^{(1)}(\vec{k}, \vec{k}')|^2 = \left| \frac{m}{2\pi\hbar^2} \int d^3\vec{r}' e^{i\vec{q} \cdot \vec{r}'} V(\vec{r}') \right|^2$$

$$\hbar\vec{q} = \vec{p}_i - \vec{p}_f, \quad \vec{q} = \vec{k} - \vec{k}'$$

For central potentials $V(\vec{r}) = V(r)$,

$$|\vec{k} - \vec{k}'| = 2k \sin \frac{\theta}{2} = q$$



$$f^{(1)}(\theta, \phi)$$

$$= -\frac{1}{2} \frac{2m}{\hbar^2} \frac{1}{iq} \int_0^\infty \frac{r^2}{r} V(r) (e^{iqr} - e^{-iqr}) dr$$

$$\int e^{iqr} \sin \theta dr = \frac{e^{iqr} - e^{-iqr}}{iq}$$

$$= -\frac{2m}{\hbar^2} \frac{1}{q} \int_0^\infty r V(r) \sin qr dr = \frac{f^{(1)}(\theta)}{q}$$

\uparrow
 $q = 2k \sin \frac{\theta}{2}$

example: $V(r) = \frac{V_0 e^{-\alpha r}}{\alpha r}$

$$f^{(1)}(\theta) = -\frac{2mV_0}{\alpha\hbar^2} \frac{1}{q^2 + \alpha^2} \left(\because \int_0^\infty r V(r) \sin qr dr = \text{Im} \int_0^\infty r V(r) e^{iqr} dr \right)$$

$$\therefore \frac{d\sigma}{d\Omega} \approx \left(\frac{2mV_0}{\hbar^2 \alpha} \right)^2 \frac{1}{[2k^2(1-\cos\theta) + \alpha^2]^2} \quad q^2 = 2k^2(1-\cos\theta)$$

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take $d \rightarrow 0$, while keeping $\frac{U_0}{\alpha}$ fixed = $ZZ'e^2$

We get: for Coulomb interaction

$$\frac{d\sigma}{d\Omega} \approx \frac{(Zm)^2 (ZZ'e^2)^2}{\hbar^4} \frac{1}{16k^4 \sin^4 \frac{\theta}{2}}$$

(Rutherford scattering cross section that was obtained in classical electrodynamics!)

Remarks:

(i) Validity of Born Approx. requires $|r' V(r')| \rightarrow 0$ as $r' \rightarrow \infty$

$$\frac{|\Psi_{sc}(0)|}{|e^{i\vec{k}\cdot\vec{r}}|} = \left| \frac{Zm}{4\pi\hbar^2} \int \frac{e^{i\vec{k}\cdot\vec{r}'}}{r'} V(\vec{r}') e^{-i\vec{k}\cdot\vec{r}'} d\vec{r}' \right| \ll 1$$

$$\uparrow$$

$$\frac{e^{i\vec{k}\cdot\vec{r}'}}{|\vec{r}-\vec{r}'|} \Big|_{r=0}$$

\therefore $|r' V(\vec{r}')| \rightarrow 0$ as $r' \rightarrow \infty$! (same as what we said in the central potential case.)

\therefore Rigorously speaking, it can not be applied to

Coulomb interaction. But it turns out that

rigorous treatment also leads to the same $\frac{d\sigma}{d\Omega}$

see Sakurai's book § 7.13 for more details.

(ii) T-matrix (transition operator)

It is often convenient to define

$$U|\psi^+\rangle = T|\phi\rangle$$

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No.

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So, the Lippmann-Schwinger equation becomes

$$V|\psi^\pm\rangle = V|\phi\rangle + V \frac{1}{E-H_0+i\epsilon} T|\phi\rangle$$

↑
T|\phi\rangle

$$\therefore T = V + V \frac{1}{E-H_0+i\epsilon} T = V + V \frac{1}{E-H_0+i\epsilon} V + \dots$$

$$\begin{aligned} \therefore f(\vec{k}', \vec{k}) &= \frac{-1}{4\pi} \frac{2m}{\hbar^2} (2\pi)^3 \int d^3\vec{p}' \frac{e^{-i\vec{k}' \cdot \vec{p}'}}{(2\pi)^{3/2}} V(\vec{p}') \langle \vec{p}' | \psi^\pm \rangle \\ &= -\frac{1}{4\pi} \frac{2m}{\hbar^2} (2\pi)^3 \int d^3\vec{p}' \langle \vec{k}' | \vec{p}' \rangle \langle \vec{p}' | V | \psi^\pm \rangle \\ &= \frac{-1}{4\pi} \frac{2m}{\hbar^2} (2\pi)^3 \int d^3\vec{p}' \langle \vec{k}' | \vec{p}' \rangle \underbrace{\langle \vec{p}' | T | \phi \rangle}_{\langle \vec{p}' | T | \vec{k} \rangle} \\ &= -\frac{1}{4\pi} \frac{2m}{\hbar^2} (2\pi)^3 \langle \vec{k}' | T | \vec{k} \rangle \end{aligned}$$

i.e. $f(\vec{k}', \vec{k})$ is the matrix element of $\frac{-1}{4\pi} \frac{2m}{\hbar^2} (2\pi)^3 T$

Higher order expansion was discussed before so we do not repeat it here.

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Relation of T matrix to S matrix

S-matrix is defined as

$$\lim_{\substack{t' \rightarrow \infty \\ t \rightarrow -\infty}} \langle f | \hat{U}(t', t) | i \rangle \equiv \langle \phi | S | \phi_i \rangle$$

$$|i\rangle = e^{-i\hat{E}_i t} |\phi_i\rangle$$

$$|f\rangle = e^{-i\hat{E}_f t} |\phi_f\rangle$$

Let us assume $H = H_0 + V$ $V =$ time-independent.This is the situation for our scattering problem in which we may set ^{when} $t < 0$, $|i\rangle = |\psi_{in}\rangle$ andfor $t > 0$, the particle comes in contact with the scattering center and get scattered!

$$\therefore U(t, 0) \Rightarrow U(t, 0) \theta(t) = e^{-i(\hat{H}_0 + \hat{V})t} \theta(t)$$

= 0 for $t < 0$

$$= \int \frac{d\omega}{2\pi} e^{i\omega t} \frac{i}{\omega - \hat{H}_0 - \hat{V} + i\epsilon}$$

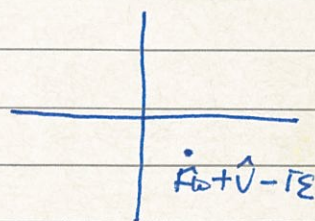
with $\epsilon \rightarrow 0^+$

$$\underbrace{\frac{i}{\omega - \hat{H}_0 - \hat{V} + i\epsilon}}_{\hat{U}(\omega)}$$

$$-i \hat{U}(\omega) = \frac{1}{\omega - \hat{H}_0 - \hat{V} + i\epsilon}$$

$$= \frac{1}{\omega - \hat{H}_0 + i\epsilon} + \frac{1}{\omega - \hat{H}_0 + i\epsilon} \uparrow \frac{1}{\omega - \hat{H}_0 + i\epsilon}$$

$$+ \frac{1}{\omega - \hat{H}_0 + i\epsilon} \uparrow \frac{1}{\omega - \hat{H}_0 + i\epsilon} \uparrow \frac{1}{\omega - \hat{H}_0 + i\epsilon} + \dots \quad \textcircled{D}$$



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$$= \frac{1}{\omega - \hat{H}_0 + i\epsilon} + \frac{1}{\omega - \hat{H}_0 + i\epsilon} \hat{T}(\omega) \frac{1}{\omega - \hat{H}_0 + i\epsilon} \quad \dots \textcircled{2}$$

where we have made use of the identity:

$$(\hat{A} - \hat{B}) \left[\frac{1}{\hat{A}} + \frac{1}{\hat{A}} \hat{B} \frac{1}{\hat{A}} + \frac{1}{\hat{A}} \hat{B} \frac{1}{\hat{A}} \hat{B} \frac{1}{\hat{A}} + \dots \right]$$

$$= (1 - \hat{B} \frac{1}{\hat{A}} + \hat{B} \frac{1}{\hat{A}} \hat{B} \frac{1}{\hat{A}} + \dots)$$

$$- \hat{B} \frac{1}{\hat{A}} - \hat{B} \frac{1}{\hat{A}} \hat{B} \frac{1}{\hat{A}} \dots = \mathbb{I}$$

$$\therefore \frac{1}{\hat{A} - \hat{B}} = \frac{1}{\hat{A}} + \frac{1}{\hat{A}} \hat{B} \frac{1}{\hat{A}} + \frac{1}{\hat{A}} \hat{B} \frac{1}{\hat{A}} \hat{B} \frac{1}{\hat{A}} + \dots$$

Note that $\hat{T} = \hat{V} + \hat{V} \frac{1}{E - \hat{H}_0 + i\epsilon} \hat{T}$

$$= \hat{V} + \hat{V} \frac{1}{E - \hat{H}_0 + i\epsilon} \hat{V} + \dots$$

Now, we can take out e^{-iEt} factor from $|i\rangle$

& $|f\rangle$ by setting $|i\rangle = e^{-iEt} |\phi_i\rangle$, $|f\rangle = e^{-iEt} |\phi_f\rangle$

$$\therefore \hat{U}(t', t) = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \hat{U}(\omega) e^{i\omega(t'-t)}$$

$$\langle f | S | i \rangle = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} e^{i(E_f t' - E_i t)} e^{-i\omega(t'-t)} \langle \phi_f | \hat{U}(\omega) | \phi_i \rangle$$

$$= \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} e^{i(E_f - \omega)t'} e^{-i(E_i - \omega)t} \frac{1}{\omega - E_i + i\epsilon} \langle \phi_f | \phi_i \rangle$$

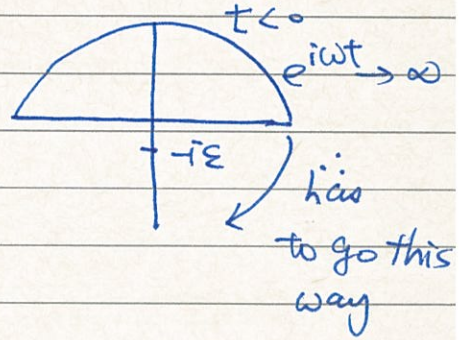
$$\textcircled{2} \quad + \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} i \frac{e^{i(E_f - \omega)t'}}{\omega - E_f + i\epsilon} \langle \phi_f | \hat{T}(\omega) | \phi_i \rangle \frac{e^{-i(E_i - \omega)t}}{\omega - E_i + i\epsilon}$$

$\dots \textcircled{3}$

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To proceed, the following identity is useful:

$$\lim_{t \rightarrow -\infty} \int_{-\infty}^{\infty} d\omega f(\omega) \frac{e^{i\omega t}}{\omega + i\varepsilon}$$



$$= \lim_{t \rightarrow -\infty} \lim_{\varepsilon \rightarrow 0^+} -2\pi i f(-i\varepsilon) e^{\varepsilon t}$$

$$= -2\pi i f(0) \quad \therefore \quad \lim_{t \rightarrow \infty} \frac{e^{i\omega t}}{\omega + i\varepsilon} = -2\pi i \delta(\omega)$$

$$\therefore \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} e^{i(E_f - \omega)t'} \quad i \frac{e^{-i(E_f - \omega)t}}{\omega - E_f + i\varepsilon} \langle \phi_f | \phi_i \rangle$$

$$\xrightarrow{t \rightarrow -\infty} \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} e^{i(E_f - \omega)t'} \quad i (-2\pi i \delta(\omega - E_f)) \langle \phi_f | \phi_i \rangle$$

$$= \frac{2\pi i}{2\pi} e^{i(E_f - E_f)t'} \langle \phi_f | \phi_i \rangle = \underline{\underline{\delta f_i}}$$

The second term in ③

$$\rightarrow \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} i \frac{e^{i(E_f - \omega)t'}}{\omega - E_f + i\varepsilon} [-2\pi i \delta(\omega - E_f)] \langle \phi_f | \hat{T}(\omega) | \phi_i \rangle$$

$$= \frac{e^{i(E_f - E_f)t'}}{E_f - E_f + i\varepsilon} \langle \phi_f | \hat{T}(E_f) | \phi_i \rangle$$

$$\xrightarrow{t' \rightarrow -\infty} \frac{e^{+i(E_f - E_f)(-t')}}{E_f - E_f + i\varepsilon} \langle \phi_f | \hat{T}(E_f) | \phi_i \rangle$$

$$\Rightarrow -2\pi i \delta(E_f - E_f) \langle \phi_f | \hat{T}(E_f) | \phi_i \rangle$$

$$\therefore \hat{S} = \hat{I} - 2\pi i \delta(E_f - E_f) \hat{T}$$

$$\Leftrightarrow S_E = 1 + 2ki a_E \quad (\text{page 6})$$

$$-\frac{i}{2} T_E(E)$$

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1.

Revisit partial wave analysis

* plane waves vs spherical waves for free particles

 (H, \vec{p}) (H, L^2, L_z) \Rightarrow see over $|\vec{k}\rangle$ $|E, \ell, m\rangle$

$$E = \frac{\hbar^2 k^2}{2m}$$

Recall that

$$\hat{k}_{||} \hat{z}, \quad e^{i\vec{k}\cdot\vec{r}} = e^{ikr\cos\theta} = \sum_{\ell=0}^{\infty} i^{\ell} (2\ell+1) j_{\ell}(kr) P_{\ell}(\cos\theta) \dots \textcircled{3}$$

This relation is actually the transformation relation between $|\vec{k}\rangle$ & $|E, \ell, m\rangle$:

$$|\vec{k}\rangle = \sum_{\ell, m} \int dE |E, \ell, m\rangle \langle E, \ell, m | \vec{k}\rangle$$

The transformation is specified by the function

$$\langle \vec{k} | E, \ell, m \rangle,$$

where $|E, \ell, m\rangle$ are normalized: $\langle E', \ell', m' | E, \ell, m \rangle = \delta(E-E') \delta_{\ell\ell'} \delta_{mm'}$

Obviously, $\langle \vec{k} | E, \ell, m \rangle$ is the solution to $H = \frac{\hat{p}^2}{2m} |\psi\rangle = E |\psi\rangle$

in the Fourier space. So it must satisfy

$$\frac{\hbar^2 k^2}{2m} \langle \vec{k} | E, \ell, m \rangle = E \langle \vec{k} | E, \ell, m \rangle \dots \textcircled{4}$$

$$\langle \vec{k} | \hat{L}^2 | E, \ell, m \rangle = \hbar^2 \ell(\ell+1) \langle \vec{k} | E, \ell, m \rangle \dots \textcircled{5}$$

$$\langle \vec{k} | \hat{L}_z | E, \ell, m \rangle = \hbar m \langle \vec{k} | E, \ell, m \rangle \dots \textcircled{6}$$

Since the angular momentum has been exactly the same form both in the real & Fourier spaces, so $\langle \vec{k} | E, \ell, m \rangle \propto Y_{\ell}^m(\hat{k})$

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$$\begin{aligned} \text{e.g. } L_z &= x p_y - y p_x = x \frac{\hbar}{i} \frac{\partial}{\partial y} - y \frac{\hbar}{i} \frac{\partial}{\partial x} = i\hbar [y \frac{\partial}{\partial x} - x \frac{\partial}{\partial y}] \\ &= i\hbar [p_y \frac{\partial}{\partial p_x} - p_x \frac{\partial}{\partial p_y}] \end{aligned}$$

$$p_x \Leftrightarrow x, \quad p_y \Leftrightarrow y.$$

Eq. (4) implies that $\langle \vec{k} | E \ell m \rangle \propto \delta(E - \frac{\hbar^2 k^2}{2m})$

\therefore We conclude

$$\langle \vec{k} | E \ell m \rangle = C \delta(E - \frac{\hbar^2 k^2}{2m}) Y_{\ell}^m(\hat{k})$$

C is fixed by the normalization $\langle E' \ell' m' | E \ell m \rangle = \delta(E-E') \delta_{\ell \ell'} \delta_{m m'}$:

$$\langle E' \ell' m' | E \ell m \rangle = \int d^3 \vec{k} \langle E' \ell' m' | \vec{k} \rangle \langle \vec{k} | E \ell m \rangle$$

$$\begin{aligned} &= |C|^2 \int k^2 dk \int d\Omega_k Y_{\ell}^m(\Omega_k) Y_{\ell'}^{m'*}(\Omega_k) \\ &\quad \times \delta(E' - \frac{\hbar^2 k^2}{2m}) \delta(E - \frac{\hbar^2 k^2}{2m}) \end{aligned}$$

$$\text{Let } E'' = \frac{\hbar^2 k^2}{2m} \quad \int k^2 dk = \frac{1}{2} \left(\frac{2m}{\hbar^2} \right)^{3/2} \sqrt{E''} dE''$$

$$\therefore \langle E' \ell' m' | E \ell m \rangle = |C|^2 \delta_{\ell \ell'} \delta_{m m'} \frac{1}{2} \left(\frac{2m}{\hbar^2} \right)^{3/2}$$

$$\times \int dE'' \sqrt{E''} \delta(E-E'') \delta(E'-E'')$$

$$= |C|^2 \frac{m k}{\hbar^2} \delta(E-E') \delta_{\ell \ell'} \delta_{m m'}$$

\therefore We can choose $C = \frac{\hbar}{\sqrt{m k}}$, thus $\langle \vec{k} | E \ell m \rangle = \frac{\hbar}{\sqrt{m k}} \delta(E - \frac{\hbar^2 k^2}{2m}) Y_{\ell}^m(\hat{k})$

L... (7)

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Combinations of eqs. (3), (2) & addition theorem

$$P_e(\cos\theta) = \frac{4\pi}{2\ell+1} \sum_m Y_\ell^m(\hat{p}) Y_\ell^{m*}(\hat{R}) \quad \text{give}$$

$$\langle \hat{p} | E \ell m \rangle = \frac{i^\ell}{k} \sqrt{\frac{2m+1}{\pi}} \sqrt{2m+1} J_\ell(kr) Y_\ell^m(\hat{r})$$

* The advantage of using the partial wave representation shows up when we measure \hat{p} from the scattering center:

incident particles coming out from the accelerator
a uniform beam



For the particles in the cylinder of radius ρ :

$$L = k\rho \approx \rho(kk)$$

\therefore If the range of the potential (r_0) is finite,

only $\ell \lesssim kr_0$ will contribute! In other words,

one needs only to consider a few partial waves.

For example, for a 100 MeV neutron incident upon

a fixed nucleus (size $\sim 10^{-15}\text{m}$), $\ell \lesssim 2$.

Only about 2 partial waves need to be considered, this

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property simplifies the whole analysis on the scattering problem! (The above argument is classical, for O.M. Res page 8.)

* Including interaction

⇒ The usage of spherical waves also simplifies the analysis when the potential is central.
~~do not need this assumption~~

In this case, $\therefore T = V + U \frac{1}{E - H_0 + i\epsilon} V + \dots$ (Scalar)

$\therefore [T, L^2] = 0, [T, L_z] = 0.$

As a result, l & m are conserved during scattering!

Technically, it means that T matrix is diagonal in $\{|Elm\rangle\}$ basis.

~~also valid in general~~
 ↓ (from Wigner-Eckart theorem)

$\langle El'm' | T | Elm \rangle = T_l(E) \delta_{l'l} \delta_{m'm}$
 ↑ depends only on l , (central potential)
 T is a scalar

For elastic scatterings, $\frac{k'_{lm}}{2m} = \frac{k_{lm}}{2m}$, therefore

$f(\vec{R}', \vec{R}) = \frac{1}{4\pi} \frac{2m}{\hbar^2} (2\pi)^3 \langle \vec{R}' | T | \vec{R} \rangle$
 $= \frac{1}{4\pi} \frac{2m}{\hbar^2} (2\pi)^3 \sum_{lm} \sum_{l'm'} \int dE \int dE' \langle \vec{R}' | El'm' \rangle \langle E'l'm' | T | Elm \rangle \langle Elm | \vec{R} \rangle$

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$$\sum_{l,m} Y_l^m(\Omega) Y_l^{m*}(\Omega') = \sum_{l=0}^{\infty} \sum_{m=-l}^l \frac{2l+1}{4\pi} P_l(\cos\theta) \quad \text{No. 5.}$$

$$= \frac{-1}{4\pi} \frac{2m}{k^2} (2\pi)^3 \frac{\hbar^2}{mk} \sum_{l,m} T_l(E) \Big|_{E=\frac{\hbar^2 k^2}{2m}} Y_l^m(\hat{R}') Y_l^{m*}(\hat{R})$$

$$= -\frac{4\pi^2}{k} \sum_{l,m} T_l(E) \Big|_{E=\frac{\hbar^2 k^2}{2m}} Y_l^m(\hat{R}') Y_l^{m*}(\hat{R})$$

$$= \sum_{l=0}^{\infty} \underbrace{\left(-\frac{\pi}{k} T_l(E)\right)}_{a_l(E)} (2l+1) P_l(\cos\theta) \quad \cos\theta = \hat{R}' \cdot \hat{R} \quad L \dots \textcircled{8}$$

$$\text{Now, } \because \psi(\vec{r}) \rightarrow \frac{1}{(2\pi)^{3/2}} \left[e^{i\vec{k} \cdot \vec{r}} + f(\vec{R}', \vec{R}) \frac{e^{ikr}}{r} \right] \dots \textcircled{9}$$

Combining eqs. $\textcircled{3}$ & $\textcircled{8}$, eq. $\textcircled{9}$ becomes

$$\psi(\vec{r}) \rightarrow \frac{1}{(2\pi)^{3/2}} \left[\sum_l (2l+1) P_l(\cos\theta) \frac{e^{ikr} - e^{-i(kr - l\pi)}}{2ikr} + \sum_l (2l+1) P_l(\cos\theta) a_l(E) \frac{e^{ikr}}{r} \right]$$

$$\left(j_l(kr) \rightarrow \frac{e^{i(kr - l\pi/2)} - e^{-i(kr - l\pi/2)}}{2ikr} \cdot i^l = e^{i\pi l/2} \right)$$

$$\rightarrow \frac{1}{(2\pi)^{3/2}} \sum_{l=0}^{\infty} (2l+1) \frac{P_l}{2ik} \left[(1 + 2i k a_l) \frac{e^{ikr}}{r} - \frac{e^{i(kr - l\pi)}}{r} \right] \dots \textcircled{10}$$

Eq. $\textcircled{10}$ makes scattering clear: scatterer changes $1 \rightarrow 1 + 2i k a_l$

i.e., it changes the outgoing waves from $\frac{e^{ikr}}{r}$ to

$$\underline{(1 + 2i a_l k) \frac{e^{ikr}}{r}} \quad ! \quad \underline{\text{Incoming waves are not changed.}}$$

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Unitarity : First, we know that l is conserved.

\therefore One can consider each l each time.

In other words, each term in (10) represents the situation: l th wave comes in and comes out.

The conservation of probability :

elastic
(inelastic, not so!)

$$\int j \cdot d\vec{s} = 0 \Rightarrow |S_l| \equiv |1 + 2iK a_l| = 1$$

(Note: $S_l \equiv \frac{\text{outgoing wave amplitude}}{\text{incoming wave amplitude}}$)

$\therefore 1 + 2iK a_l = e^{i\delta_l}$ $S_l \equiv \text{phase shift.}$ $i \equiv (-2\pi i T_l(E))$

$$a_l = \frac{e^{i\delta_l} - 1}{2iK} = \frac{e^{i\delta_l} \sin \delta_l}{K} \quad \text{i.e. } R(r) \rightarrow \frac{1}{r} \sin(kr - \frac{\delta_l}{2} + \delta_l)$$

$$\therefore f(\theta) = \sum_{l=0}^{\infty} \frac{(2l+1)}{K} \underbrace{e^{i\delta_l} \sin \delta_l P_l(\cos \theta)}_{(2l+1) a_l(r)} \dots (11)$$

total cross section $\equiv \sigma_t = \int |f|^2 d\Omega$

$$= \frac{4\pi}{K^2} \sum_{l=0}^{\infty} (2l+1) \sin^2 \delta_l = 4\pi \sum_{l=0}^{\infty} (2l+1) |a_l|^2 \dots (12)$$

$$\int P_l(\cos \theta) P_{l'}(\cos \theta) d\cos \theta = \frac{2}{2l+1} \delta_{ll'}$$

"unitary limit"

$$\max_{\delta_l} \sigma_l = \frac{4\pi}{K^2} (2l+1)$$

$\therefore \sigma_t$ depends only on phase shifts.

From (11) & (12) \Rightarrow optical theorem $\sigma_t = \frac{4\pi}{K} \text{Im}(f(0))$

"unitary $\Rightarrow \delta_l = \text{real}$ "

$$\therefore \sin^2 \delta_l \leq 1$$

Determination of phase shifts

In terms of δ_l , eq. (10) becomes

$$\psi(r, \theta) \rightarrow \frac{1}{(2\pi)^{3/2}} \sum_{l=0}^{\infty} (2l+1) i^l P_l \left[\frac{e^{2i\delta_l} e^{i(kr - \frac{\pi}{2})} - e^{-i(kr - \frac{\pi}{2})}}{2iKr} \right] \dots (13)$$

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provide a cut off of $l!$
 \downarrow gate

for $l \gg \sqrt{KR} = l_{max}$

$$\therefore \tan \delta_l = \frac{KR j_l'(KR) - \beta_l j_l(KR)}{KR n_l'(KR) - \beta_l n_l(KR)}$$

$$j_l(KR) \sim \frac{1}{(2l+1)!!} (KR)^l$$

$$n_l(KR) \sim (2l+1)!! (KR)^{l-1}$$

$$\beta_l \sim -\frac{l-\beta_l}{l+1+\beta_l} \frac{(KR)^{2l+1}}{(2l+1)!!(2l)!!} \ll 1$$

\therefore To determine δ_l , we need only to solve

the 1D Schrödinger equation (usually numerically) up

to $r=R$ to get $\beta_l(R)$ (by continuity, $\beta_l(R^-) = \beta_l(R^+)$):

$$\therefore \delta_l = \frac{4\pi}{k^2} \sum_{l=0}^{l_{max}}$$

$$\frac{d^2}{dr^2} [rA_l] + \left[k^2 - \frac{2m}{\hbar^2} V - \frac{l(l+1)}{r^2} \right] (rA_l) = 0 \quad (2l+1) \sin^2 \delta_l$$

$$rA_l |_{r=0} = 0.$$

example: Hard sphere



$$A_l(R) = 0 \quad e^{i\delta_l} [\cos \delta_l j_l(KR) - \sin \delta_l n_l(KR)] = 0$$

$$\therefore \tan \delta_l = \frac{j_l(KR)}{n_l(KR)}$$

$$\therefore j_0(x) = \sin x/x, \quad n_0(x) = -\cos x/x$$

$$\therefore \tan \delta_0 = -\tan(KR), \quad \delta_0 = -KR$$

$$\begin{aligned} \text{i.e. } A_0(r) &= e^{i\delta_0} [\cos \delta_0 j_0(Kr) - \sin \delta_0 n_0(Kr)] \\ &= e^{i\delta_0} \frac{\sin(Kr + \delta_0)}{Kr} = e^{-iKR} \frac{\sin K(R-r)}{Kr} \end{aligned}$$

\therefore The hard sphere has just pushed out the wave function, forcing it to start the oscillation at $r=R$!

Note that $\therefore j_l(x) \rightarrow x^l / (2l+1)!!$, $n_l(x) \rightarrow -x^{-(2l+1)} / (2l)!!$, $x \rightarrow 0$
 $\therefore \beta_l \sim -(KR)^{2l+1}$ at low energies, \therefore Any small l get phase shifted, i.e. get scattered!

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Recall that

$$j_\ell(kr) \rightarrow \frac{\sin(kr - \frac{\ell\pi}{2})}{kr}$$

as $r \rightarrow \infty$

$$n_\ell(kr) \rightarrow \frac{-\cos(kr - \frac{\ell\pi}{2})}{kr}$$

$$\therefore j_\ell(kr) + i n_\ell(kr) \rightarrow \frac{1}{ikr} e^{i(kr - \frac{\ell\pi}{2})}$$

$$j_\ell(kr) - i n_\ell(kr) \rightarrow \frac{-1}{ikr} e^{-i(kr - \frac{\ell\pi}{2})}$$

$$[\dots] \text{ in } (13) = e^{i\delta_\ell} \frac{j_\ell(kr) + i n_\ell(kr)}{2} + \frac{j_\ell(kr) - i n_\ell(kr)}{2}$$

$$= e^{i\delta_\ell} [\cos \delta_\ell j_\ell(kr) - \sin \delta_\ell n_\ell(kr)]$$

$$\therefore \psi(r) \rightarrow \frac{1}{(2\pi)^{3/2}} \sum_{\ell=0}^{\infty} i^\ell (2\ell+1) A_\ell(r) P_\ell(\cos\theta)$$

$$\rightarrow \frac{e^{i\delta_\ell}}{kr} \sin(kr - \frac{\ell\pi}{2} + \delta_\ell)$$

$$A_\ell(r) \equiv e^{i\delta_\ell} [\cos \delta_\ell j_\ell(kr) - \sin \delta_\ell n_\ell(kr)] \dots (14)$$

Which satisfies
$$\left[\frac{-\hbar^2}{2m} \frac{d^2}{dr^2} + \frac{\ell(\ell+1)\hbar^2}{2mr^2} \right] (rA_\ell) = E (rA_\ell) \quad E = \frac{\hbar^2 k^2}{2m}$$

This is nothing but the statement: $r \rightarrow \infty$ $\psi \rightarrow$ solutions

of free particles! (We have actually looked at this problem

via this point of view when we discussed central potentials.)

For a finite range potential ($r < R$), eq. (14) is

valid only for $r \geq R$. To extract δ_ℓ , one forms

a dimensionless parameter

$$\beta_\ell \equiv \left. \frac{r}{A_\ell} \frac{dA_\ell}{dr} \right|_{r=R} = KR \frac{j'_\ell(KR) \cos \delta_\ell - n'_\ell(KR) \sin \delta_\ell}{j_\ell(KR) \cos \delta_\ell - n_\ell(KR) \sin \delta_\ell}$$

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Zero-energy scattering & scattering length

$$\sigma = \sum_l \frac{4\pi}{k^2} (2l+1) \sin^2 \delta_l$$

Consider the case $U=0$ for $r \geq R$

$$\rightarrow 4\pi \left(\frac{f_0}{k}\right)^2$$

and $E \rightarrow 0$

$$= 4\pi a^2$$

in $E \rightarrow 0$

then $\frac{-\hbar^2}{2m} \frac{d^2}{dr^2} (rR_l) = 0$

u_l

for $l=0, r > R$

($E \rightarrow 0$ only $\delta_{l=0}$ survives
see below!)

$\therefore U_0 = \text{const. } (r-a)$

$a \equiv$ scattering length

This can be also view^{ed} as the limit of

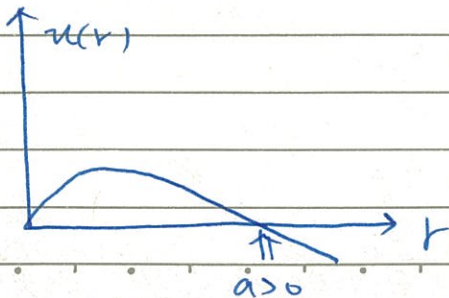
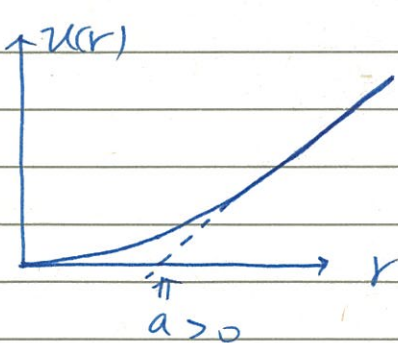
$$\lim_{k \rightarrow 0} \sin(kR + \delta_0) = \lim_{k \rightarrow 0} \sin k \left(r + \frac{\delta_0}{k} \right)$$

U_0

\uparrow
 $-a$

$\delta_0 \sim (kR)^{2l+1}$

a can be both positive or negative:



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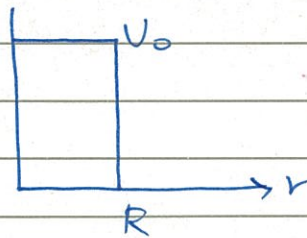
2,

a can differ from R by orders of magnitude.

For hard spheres, $a=R$ ($\because u = \sin k(r-R) \quad r \geq R$)

There is no general relation between attractive/repulsive and sign of a :

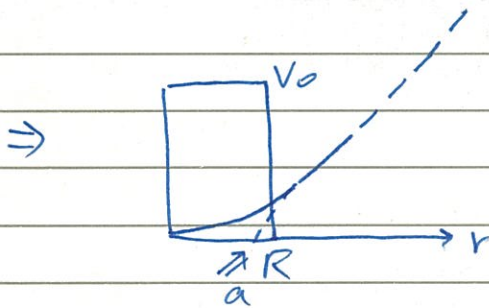
repulsive:



$E \rightarrow 0$ \uparrow ($u(r)$ is $\sinh kr$
for $0 \leq r \leq R$ $\frac{\hbar^2 k^2}{2m} = V_0$)
($u(0)=0$)

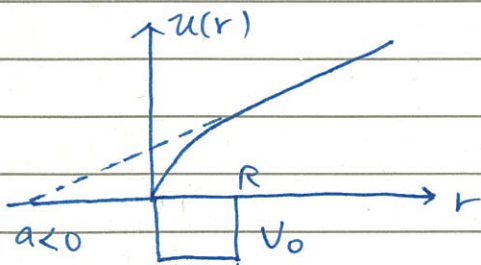
$$\therefore u' = \cosh kr$$

$$u'(R) = \cosh kR > 0$$



$$a > 0$$

attractive: $V_0 < 0$



$|V_0|$ small enough

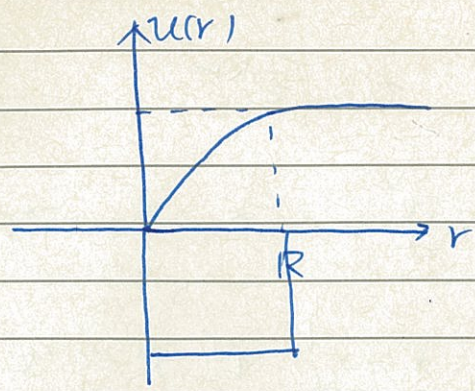
$$u(r) \sim \sin kr \quad k = \sqrt{\frac{2m|V_0|}{\hbar^2}} \text{ is small}$$

\nearrow Sin

so that $u'(R) > 0$!

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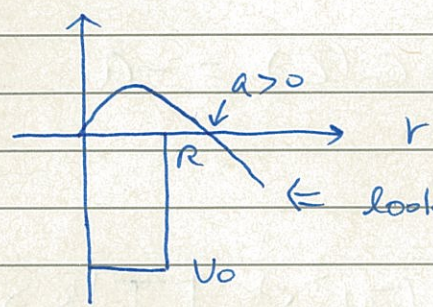
intermediate V_0



$a = \infty$

e.g. $KR = \frac{\pi}{2}$ $|V_0| = \frac{\hbar^2}{2m} \left(\frac{1}{R}\right)^2 \cdot \frac{\pi^2}{4}$

large V_0



looks like exponential decay if

bound state may form if $a \gg R$ \Rightarrow see over

Cross section : $\therefore a_l = \frac{e^{i\delta_l} \sin \delta_l}{k} = \frac{\sin \delta_l}{k(\cos \delta_l - i \sin \delta_l)}$
 $= \frac{1}{k \cot \delta_l - i k}$

$\therefore \lim_{k \rightarrow 0} a_l(k) = \lim_{k \rightarrow 0} \frac{1}{k \cot \delta_l} \dots \textcircled{2}$

but we have seen $\delta_l \sim k^{2l+1} \therefore a_l(k) \sim \frac{\sin \delta_l}{k} \sim 0$

for $l \neq 0$ $k \rightarrow 0$

\therefore only $a_0(k)$ survives (s-wave, $\delta_0(k)$)

$\therefore \sigma_{tot}(k \rightarrow 0) = 4\pi \left(\lim_{k \rightarrow 0} a_0(k) \right)^2$

quick way
 $\frac{\sin \delta_l}{k} \sim \frac{\delta_l}{k} = (-a)$
 $\delta_l \rightarrow 0 \quad e^{i\delta_l} \rightarrow 1$

$$a \gg R \quad r > R \quad u \sim e^{-kr}$$

$$\therefore \frac{u'}{u} = -k = \frac{1}{r-a} \Big|_{r=R} \quad (\text{page 3})$$

$$\therefore k \approx \frac{1}{a} \quad \text{if } R \ll a$$

$$\therefore E = -\frac{\hbar^2 k^2}{2m} = \frac{-\hbar^2}{2m} \left(\frac{1}{a}\right)^2$$

$$f(k', k) \rightarrow a_0(E) = -a$$

↑
eg 8

$$= \frac{-1}{4\pi} \frac{2m}{\hbar^2} (2\pi)^3 \langle k' | T | k \rangle$$

||

V_0 Born-Approx
 $\delta(r)$

$$= \frac{-1}{4\pi} \frac{2m}{\hbar^2} V_0$$

$$V = \frac{2\pi \hbar^2}{m} \delta(r)$$

$$\therefore V_0 = \frac{2\pi \hbar^2}{m}$$

↑
pseudo potential

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Now, from (1)

$$\frac{u_0'}{u_0} = k \cot k \left(r + \frac{\delta_0}{k} \right)$$

$$\xrightarrow[k \rightarrow 0]{r=0} k \cot \delta_0 \text{ should } \rightarrow \left. \frac{1}{r-a} \right|_{r=0} \quad \leftarrow \text{over}$$

$$\therefore \lim_{k \rightarrow 0} \frac{1}{k \cot \delta_0} = -a = \lim_{k \rightarrow 0} a_0(k)$$

→ pseudo potential

$$U(r) = \frac{4\pi a^2 \hbar^2}{m} \delta(r)$$

(Fermi, 1936)

$$\therefore \delta_0(k \rightarrow 0) = \frac{4\pi a^2}{4!} !$$

ie. the intercept of the zero energy wavefunction $u(r)$ determines $\delta_0(k \rightarrow 0)$!

Friedel Sum rule

Consider putting the scattering event inside a box L^3 , the scattering potential is at the center. A good example is the electron inside the metal, scattered by impurities.

In the absence of impurities,

$\therefore R_{kl}(r) \sim \frac{1}{r} \sin(kr - \frac{1}{2}l\pi)$ has to vanish at $r=L$

$$\therefore k_{nl} = \frac{n\pi}{L} + \frac{\delta_{nl}}{2L} \therefore \frac{dk}{dn} = \frac{\pi}{L}, \quad \frac{dn}{dk} = \frac{L}{\pi}$$

= density of states for fixed l

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In the presence of impurities

$$k_{nl} = \frac{n\pi}{L} + \frac{\rho_n}{2L} - \frac{\delta\epsilon(k_{nl})}{L}$$

$$\therefore 1 = \frac{\pi}{L} \frac{dn}{dk} - \frac{1}{L} \frac{d\delta\epsilon}{dk}$$

$$\therefore \frac{dn}{dk} = \frac{L}{\pi} + \frac{d\delta\epsilon}{\pi dk}$$

Obviously, the extra states are due to impurities.

$$\frac{d\delta n}{dk} = \frac{1}{\pi} \frac{d\delta\epsilon}{dk}$$

$\therefore \frac{dN}{dk} \equiv$ total change in the density of states

$$= \frac{d}{dk} \sum_{\substack{\ell, m_\ell \\ m_s}} \frac{\delta\epsilon}{\pi}$$

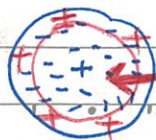
In the case of metal, electrons are filled up to k_F

$$\Rightarrow \int_0^{k_F} \frac{dN}{dk} dk = \text{total change of electron \#}$$

$$= \sum_{\substack{m_s, m_\ell, \ell}} \frac{\delta\epsilon(k_F)}{\pi}$$

To maintain charge neutrality, total change of electron # = charge on the impurity = Z

$$\therefore Z = \sum_{\substack{m_s, m_\ell \\ \ell}} \frac{\delta\epsilon(k_F)}{\pi} \quad \Leftarrow \text{Friedel Sum Rule}$$



neutral even though there will be + charge (due to exposed ions) on the outer surface!

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Analytic properties of $S_{\ell=0}$ & Resonance scattering

$$\frac{\text{The ratio of outgoing}}{\text{incoming wave}} = S_{\ell} = 1 - 2\pi i T_{\ell}(E)$$

In other words, for $\ell=0$,

$$R \sim S_{\ell=0}(k) \frac{e^{ikr}}{r} - \frac{e^{-ikr}}{r}$$

$$\text{Where } E = \frac{\hbar^2 k^2}{2m} \quad (\text{or } E - U(r \rightarrow \infty) = \frac{\hbar^2 k^2}{2m})$$

$$\text{Now, suppose } E < 0, \text{ i.e. } k \rightarrow i\kappa \quad \frac{\hbar^2 k^2}{2m} = -E$$

and it corresponds to a bound state, one expects:

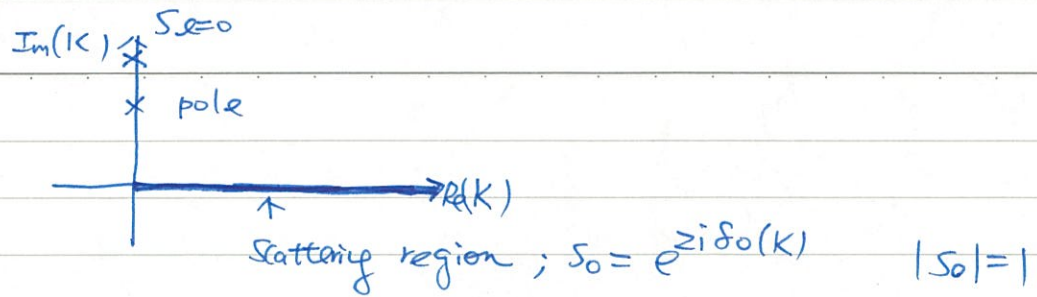
$$\left\{ \begin{array}{l} \text{no incident wave} \\ \text{only} \end{array} \right. \therefore \frac{e^{-\kappa r}}{r} \text{ is present.}$$

$$\text{In other words, } \frac{\text{outgoing}}{\text{incoming}} = \infty \quad \therefore \lim_{k \rightarrow 0} S_{\ell=0}(k) = \infty$$

Therefore, we expect that if we consider $S_{\ell=0}(k)$ as a complex number, $S_{\ell=0}(k)$ has a pole at $i\kappa$!

It can be shown that such poles are simple poles (for $\ell=0$ & $\ell \neq 0$!) Ref. Scattering theory, Aleksei G. Sitenko, Springer-Verlag, 1991

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Note that $\therefore \lim_{k \rightarrow 0} \frac{1}{k \cot \delta_0} = -a$

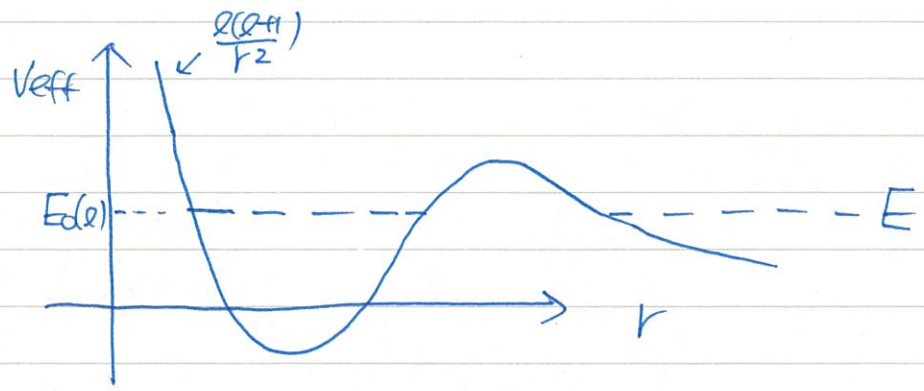
$\therefore \delta_0(k \rightarrow 0) = 0, \pm\pi, \dots \therefore S_0(k=0) = 1$

Resonance Scattering

In some cases, when $E > 0$, ^{even though} the system does not have a bound state, it may happen that

the effective $V_{eff} = V(r) + \frac{\hbar^2}{2m} \frac{l(l+1)}{r^2}$ has

a potential well that may trap the particle in a "quasi-bound" state. (e.g. Feshbach resonance)



Such a state has a finite lifetime due to quantum mechanical tunneling.

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Since $\sigma = \sigma_r = \frac{4\pi}{k^2} (2l+1) \sin^2 \delta_l$ for "l-wave", one

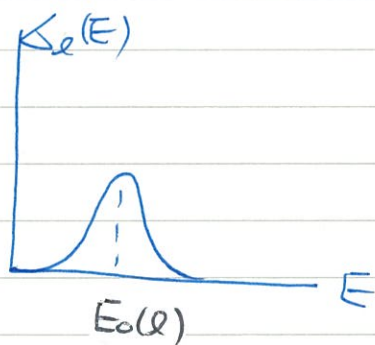
expects that as we vary E through $E_0(l)$, σ_r

will go through a max. right at $E_0(l)$!

If only one quasi-bound state is present,

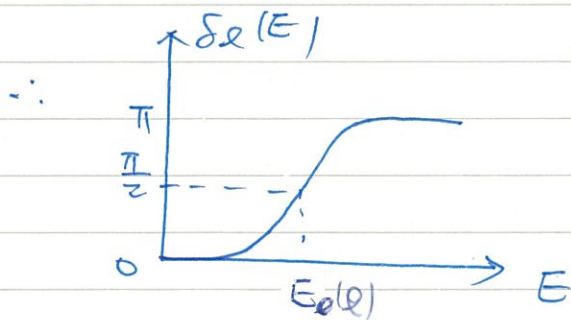
& for $E \ll E_0(l)$, and $E \gg E_0(l)$, there is

no scattering. We have



the phase shifts at two limits $E \ll E_0(l)$

& $E \gg E_0(l)$ are 0 & π ($S_l = e^{2i\delta_l}$)

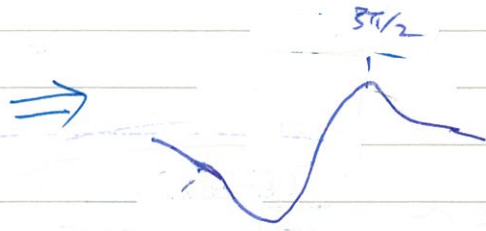
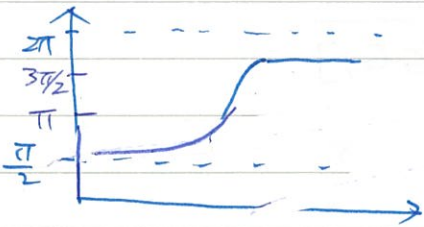
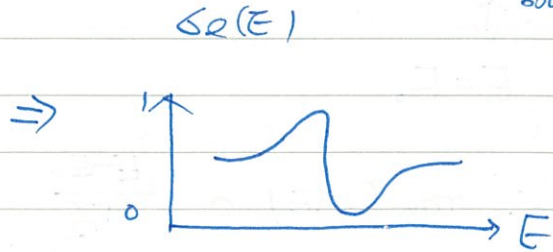
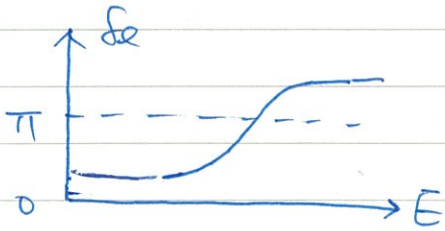


$$\delta_l(E) = \tan^{-1} \left(\frac{T/2}{E_0 - E} \right)$$

\therefore At $E_0(l)$, $\sigma_r = \frac{4\pi}{k^2} (2l+1)$ (unitary limit)

reaches max !!

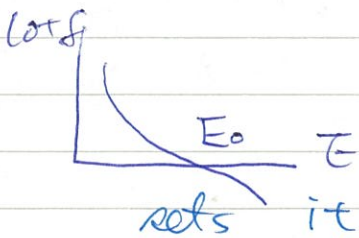
There are also other ways to go through the resonant energy $E_0(E)$: (due to background; other quasi-bound states)



We will concentrate on the 1st case (others see over)

In this case, $\because \delta_l(E_0) = \pi/2 \therefore \cot \delta_l |_{E=E_0} = 0$

$$\therefore \cot \delta_l(E) \approx + \left. \frac{d \cot \delta_l}{dE} \right|_{E=E_0} (E-E_0) \Rightarrow \delta_l \approx \tan^{-1} \left(\frac{\Gamma/2}{E_0-E} \right)$$



$\because \left. \frac{d \cot \delta_l}{dE} \right|_{E=E_0} < 0 \therefore$ Usually, one

sets it to $\frac{d \cot \delta_l}{dE_0} = -\frac{2}{\Gamma}$

$$\therefore a_l = \frac{e^{i\delta_l} \sin \delta_l}{k} = \frac{1}{k \cot \delta_l - i k} = -\frac{\Gamma/2}{k[(E-E_0) + i\Gamma/2]}$$

$$\therefore \Delta E = \frac{4\pi a_0^2}{(2\mu\hbar)} = \frac{4\pi}{k^2} \frac{(2\mu\hbar)(\Gamma/2)^2}{(E-E_0)^2 + \Gamma^2/4}$$

This is known as the Breit-Wigner formula.

$$\Delta E = \Gamma/2, \quad \therefore \Delta t \sim \frac{\hbar}{2 \Delta E} = \frac{\hbar}{\Gamma}$$

lifetime of quasi-bound state!

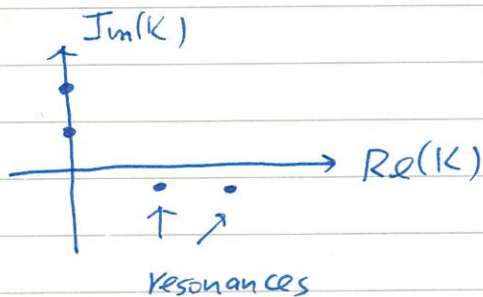
Note that $\therefore \delta \approx \tan^{-1} \frac{\Gamma/2}{E_0 - E}$

$$\therefore S_e(k) = e^{2i\delta} = \frac{e^{i\delta}}{e^{-i\delta}}$$

$$= \frac{1 + i \tan \delta}{1 - i \tan \delta} = \frac{E - E_0 - i\Gamma/2}{E - E_0 + i\Gamma/2}$$

\therefore resonance also corresponds to a pole

at $E_0 - i\Gamma/2$. $\left(e^{-iEt/\hbar} = e^{-iE_0 t/\hbar} e^{-\frac{\Gamma t}{2\hbar}} \right)$



$$\Delta t = \frac{\hbar}{\Gamma}$$

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Fool's day!

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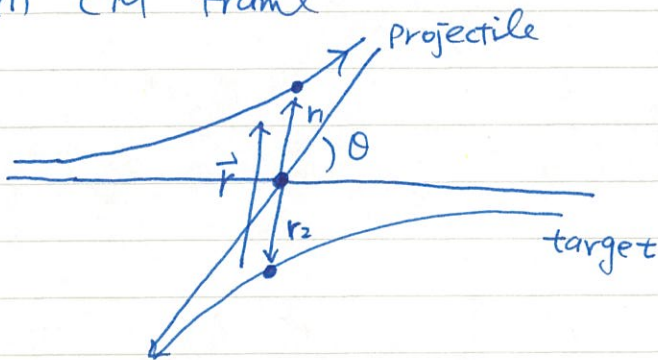
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Two-particle scattering

The above formulation applies to the case when the scattering center is fixed.

When we consider scattering for two particles, there are two useful frames that we can look at this problem.

(i) CM frame



$$\begin{cases} \vec{v}_1 = \vec{v}_{CM} + \frac{m_2}{M} \vec{v} \\ \vec{v}_2 = \vec{v}_{CM} - \frac{m_1}{M} \vec{v} \end{cases}$$

write \vec{v}_1 & \vec{v}_2 in terms of \vec{v} & \vec{v}_{CM}

$$\vec{v}_{CM} = \frac{m_1 \vec{v}_1 + m_2 \vec{v}_2}{M}$$

$$\psi_{inc} = e^{i\vec{k}_1 \cdot \vec{r}_1} e^{i\vec{k}_2 \cdot \vec{r}_2} = e^{i\vec{K} \cdot \vec{R}_{CM}} e^{i\vec{k} \cdot \vec{r}} \quad \vec{r} = \vec{r}_1 - \vec{r}_2$$

$$\vec{K} = \vec{k}_1 + \vec{k}_2 \quad \vec{k} = \frac{1}{M} (m_2 \vec{k}_1 - m_1 \vec{k}_2) \quad \text{the } \vec{k} = \text{momentum of reduced mass } \frac{1}{\mu} = \frac{1}{m_1} + \frac{1}{m_2}$$

$$\text{the } \vec{k} = \frac{1}{M} (m_2 \text{the } \vec{k}_1 - m_1 \text{the } \vec{k}_2) \quad \text{fictitious particle.}$$

$$= \frac{1}{M} (m_2 m_1) (\vec{v}_1 - \vec{v}_2) = \mu \vec{v}_{rel}$$

$e^{i\vec{k} \cdot \vec{R}_{CM}}$ is not changed during scattering, but

$$e^{i\vec{k} \cdot \vec{r}} \rightarrow e^{i\vec{k} \cdot \vec{r}} + \frac{f(\theta, \phi)}{r} e^{i\vec{k} \cdot \vec{r}}$$

$$\therefore \psi(r_1, r_2) \rightarrow e^{i\vec{K} \cdot \vec{R}_{CM}} \left[e^{i\vec{k} \cdot \vec{r}} + \frac{f(\theta, \phi)}{r} e^{i\vec{k} \cdot \vec{r}} \right]$$

In the CM. frame, one has $\vec{K} = 0$, so $e^{i\vec{K} \cdot \vec{R}_{CM}} = 1$

can be dropped out. Since $\vec{r}_1 \parallel \vec{r}_2$, to take for the

projectiles scattered into $d\Omega$ is the same as that for the fictitious particles scattered into $d\Omega$!

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i.e.

projectiles / per sec through $d\Omega$

$$= |f(\theta, \phi)|^2 \frac{\hbar k}{u} d\Omega \cdot \frac{1}{V} \quad (\text{where we have used}$$

the result from fixed scattering center problem.)

In the CM frame, $\vec{k} = \frac{1}{M} (m_2 \vec{k}_1 - m_1 \vec{k}_2) = \vec{k}_1 = -\vec{k}_2$.

$\vec{k}_2 = -\vec{k}_1$

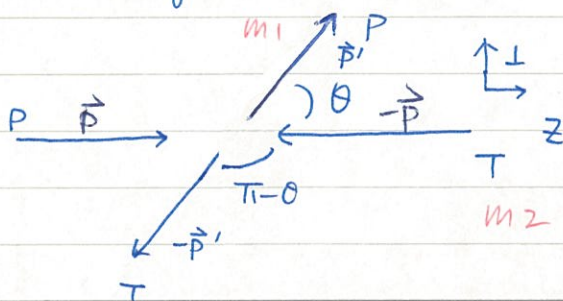
$$\begin{aligned} \text{relative incident flux} &= \frac{\hbar k_1}{m_1} \frac{1}{V} - \frac{\hbar k_2}{m_2} \frac{1}{V} \\ &= \frac{\hbar k}{V} \left(\frac{1}{m_1} + \frac{1}{m_2} \right) \\ &= \frac{\hbar k}{u} \cdot \frac{1}{V} \end{aligned}$$

\therefore The whole problem is the same as the fixed scattering center problem except that one has to consider the fictitious particle with a reduced mass μ .

(ii) Lab frame

In the lab frame, the target is at rest initially.

The passage to the lab frame from the CM frame is straight forward:



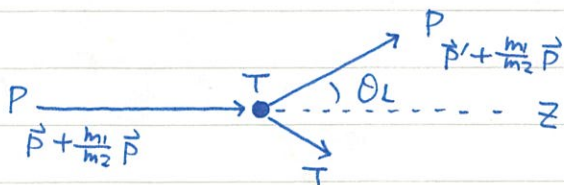
$$\text{C.M. } \tan \theta = \frac{P_{\perp}'}{P_z'}$$

$$\tan \phi = \frac{P_y'}{P_x'}$$

$$|P'| = |P| = P$$

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The same scattering event occurs at different angles in the lab frame.



In the lab frame, T is at rest initially, so we shift the momentum of projectiles by $\vec{p}' + \frac{m_1}{m_2} \vec{p}$

$$\therefore \tan \theta_L = \frac{p_L'}{p_z' + \frac{m_1}{m_2} p} = \frac{p_L'/p'}{\frac{p_z'}{p'} + \frac{m_1}{m_2}} = \frac{\sin \theta}{\cos \theta + m_1/m_2}$$

$$|p'| = |p|$$

$$\tan \phi_L = p_y'/p_x' = \tan \phi, \quad \phi_L = \phi$$

$$\therefore \frac{d\phi}{d\Omega_L} d\Omega_L = \frac{d\phi}{d\Omega} d\Omega \quad \dots \quad (15)$$

$$\therefore \frac{d\phi}{d\Omega_L} = \frac{d\phi}{d\Omega} \frac{d\Omega}{d\Omega_L} = \frac{d\phi}{d\Omega} \frac{\sin \theta d\theta d\phi}{\sin \theta_L d\theta_L d\phi}$$

By expressing θ in terms of θ_L , we get

$$\frac{d\phi}{d\Omega_L}(\theta_L)$$

For example, $m_1 = m_2$, $\therefore \frac{\sin \theta}{\cos \theta + 1} = \tan(\theta/2) \therefore \theta_L = \theta/2$

$$\frac{\sin \theta d\theta}{\sin \theta_L d\theta_L} = \frac{\sin 2\theta_L}{\sin \theta_L} = 4 \cos \theta_L \therefore \frac{d\phi}{d\Omega_L} = 4 \cos \theta_L \frac{d\phi}{d\Omega} (2\theta_L)$$

Note that because of eq. (15), σ_{total} is the same in both frames.

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Scattering of identical particles

In the above discussion, the calculated cross section is only for the projectiles.

If the detector is sensible only to the projectile,

$\frac{d\sigma}{d\Omega}$ is the cross section one measures:

$$\frac{d\sigma}{d\Omega} = |f(\theta, \phi)|^2$$

If the detector just counts # of particles that enter into it, we have

$$\frac{d\sigma}{d\Omega} = |f(\theta)|^2 + |f(\pi - \theta)|^2 \quad \text{for}$$

non-identical particles.

For identical particles, one needs to consider receiving target particles automatically no matter what kind of sensor one uses.

For spin-zero bosons, we need to symmetrize the total wave function.

$$\therefore \psi_{\text{init.}} = e^{i\vec{R} \cdot \vec{R}_{CM}} (e^{i\vec{k} \cdot \vec{r}} + e^{-i\vec{k} \cdot \vec{r}})$$

$$\psi(\vec{r} \rightarrow \infty) = \underbrace{e^{i\vec{k} \cdot \vec{r}} + e^{-i\vec{k} \cdot \vec{r}}}_{\text{initial}} + [f(\theta, \phi) + f(\pi - \theta, \phi + \pi)] \frac{e^{ikr}}{r}$$

actually, no need for symmetrization as initially they are well separated and are

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$$\therefore \frac{d\sigma}{d\Omega} = |f(\theta, \phi) + f(\pi - \theta, \phi + \pi)|^2 \quad \begin{matrix} \frac{\hbar \vec{k}_1}{m} \frac{1}{v} - \frac{\hbar \vec{k}_2}{m} \frac{1}{v} \\ \uparrow \end{matrix}$$

(note that the incident flux is still the same because initially projectiles & targets are well separated, and thus distinguishable!)

$$= \underbrace{|f(\theta, \phi)|^2 + |f(\pi - \theta, \phi + \pi)|^2}_{\text{distinguishable case}} + 2\text{Re}[f(\theta, \phi)f^*(\pi - \theta, \phi + \pi)]$$

* One sees that in addition to $|f(\theta, \phi)|^2 + |f(\pi - \theta, \phi + \pi)|^2$, one gets interference term $2\text{Re}[f(\theta, \phi)f^*(\pi - \theta, \phi + \pi)]$!

If $f = f(\theta)$, then at $\theta = \pi/2$, the cross section $= 4|f(\pi/2)|^2$ is enhanced by 2 times!

* To find the total cross section σ , we integrate over only 2π not 4π :

$$\sigma = \int_0^{\frac{\pi}{2}} d\theta \int_0^{2\pi} d\phi \sin\theta \frac{d\sigma}{d\Omega}$$

This is not to double count each event twice!

For Fermions, the total wave function needs to be antisymmetrized.

For example, consider scattering of two identical electrons.

If the incident beam is unpolarized, one has $\frac{1}{4}$ contribution coming from singlet, $\frac{3}{4}$ from triplet.

Therefore,

$$\frac{d\delta}{dJ_2} = \frac{1}{4} |f(\theta, \phi) + f(\pi - \theta, \phi + \pi)|^2$$

↑
Symmetric in space part
singlet

$$+ \frac{3}{4} |f(\theta, \phi) - f(\pi - \theta, \phi + \pi)|^2$$

↑
antisymmetric in space part
triplet.